



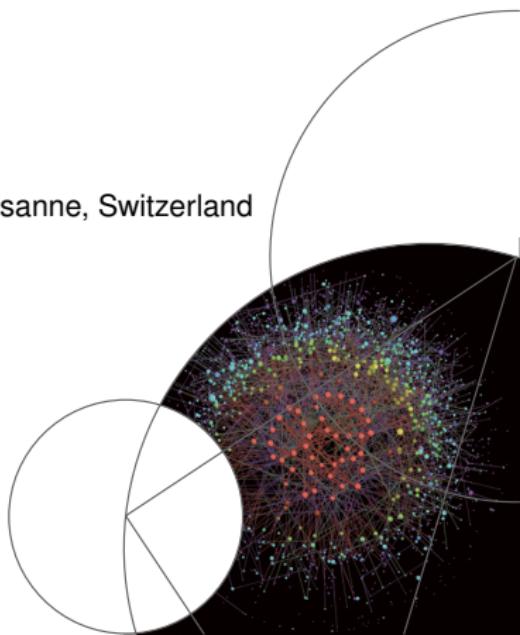
Entropic selection of concepts in networks of similarity between documents

Alessio Cardillo

Laboratory for Statistical Biophysics (LBS)

École polytechnique fédérale de Lausanne (EPFL), Lausanne, Switzerland

<http://bifi.es/~cardillo/>



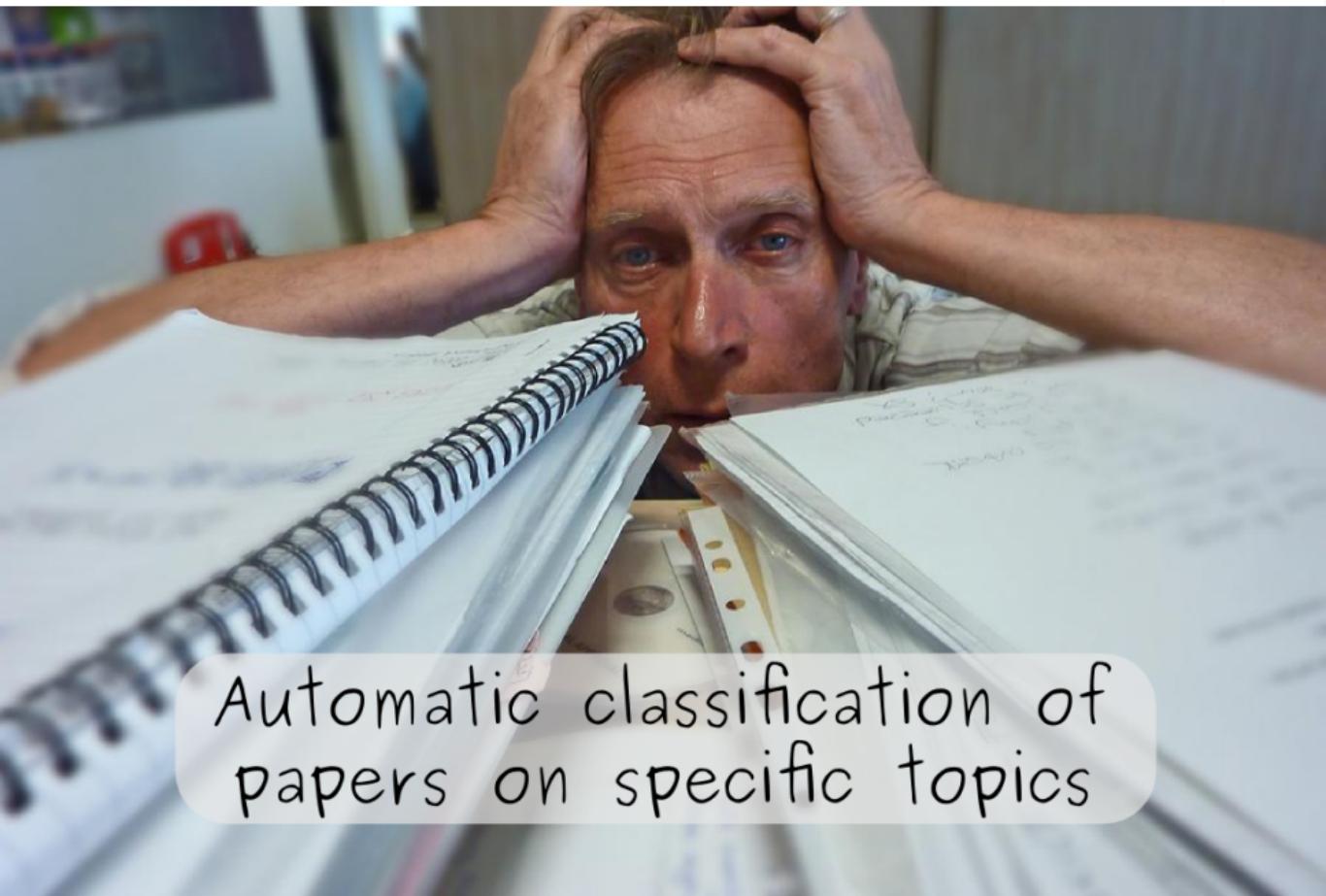
Acknowledgements

Main Collaborators

- Paolo De Los Rios (EPFL)
- Andrea Martini (EPFL)

Other Collaborators

- Vasyl Palchykov, Valerio Gemmetto, Diego Garlaschelli
(Leiden - NL)
- ScienceWISE Team (Lausanne)



Automatic classification of
papers on specific topics



Recent ontology graph



Recently bookmarked papers

Properties of a possible class of particles
astro-ph/9505117 Luis Gonzalez-Mestres

The apparent Lorentz invariance of the laws of physi...

Introduction to the Standard Model and E
0901.0241 Paul Langacker

A concise introduction is given to the standard model. Including the structure of the QCD and electroweak Lagrangians, spontaneous symmetry breaking, experimental tests, and problems.

[Standard Model](#) [Quantum chromodynamics](#)
[Weak interaction](#) ...

<http://sciencewise.info>

Outline

- Introduction on similarity networks.
- ★ Filtering of weighted networks.
- ★ Entropic filtering of concepts.
- ★ Results.
 - Take home messages
 - Questions

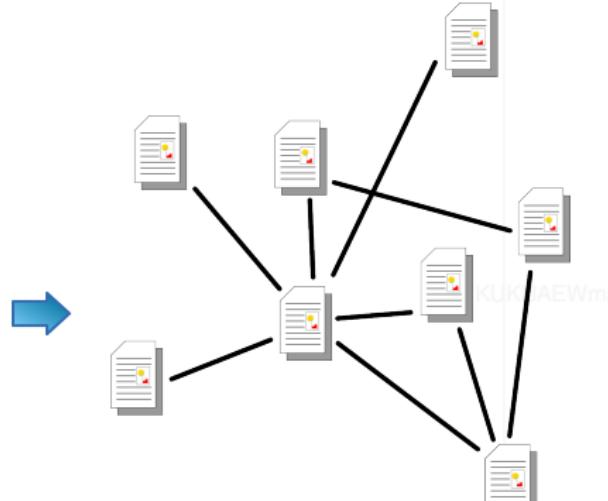
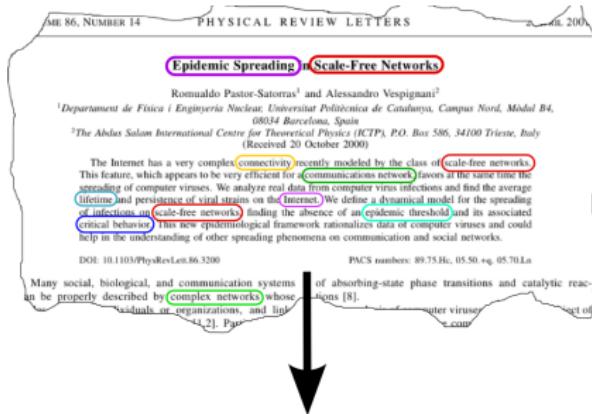
IAEWm

Section 1

Similarity networks

KUKUAEWm

Networks of similarity between papers



Epidemics spreading

Scale-free networks

Connectivity

Communications nets

Lifetime

...

Similarity network

...

TF-IDF and similarity

$$TF-IDF_{\alpha c} = u_{\alpha c} = tf_{\alpha c} \log \left(\frac{N}{df_c} \right).$$

α

| | | | | | | |
|---|---|---|---|---|---|---|
| 1 | 1 | 1 | 0 | 1 | 0 | 0 |
|---|---|---|---|---|---|---|

$c_1 \ c_2 \ c_3 \ c_4 \ c_5 \ c_6 \ c_7$

KUKUAEWm

TF-IDF and similarity

Edge weight

$$w_{\alpha\beta} = \frac{\vec{u}_\alpha \cdot \vec{u}_\beta}{\|\vec{u}_\alpha\| \|\vec{u}_\beta\|},$$

$$w_{\alpha\beta} \in [0, 1],$$

| | | | | | | | |
|----------|-------|-------|-------|-------|-------|-------|-------|
| α | 2 | 43 | 0 | 18 | 0 | 11 | 27 |
| β | 13 | 5 | 9 | 0 | 30 | 0 | 0 |
| | C_1 | C_2 | C_3 | C_4 | C_5 | C_6 | C_7 |

TF-IDF and similarity

Edge weight

$$w_{\alpha\beta} = \frac{\vec{u}_\alpha \cdot \vec{u}_\beta}{\|\vec{u}_\alpha\| \|\vec{u}_\beta\|},$$

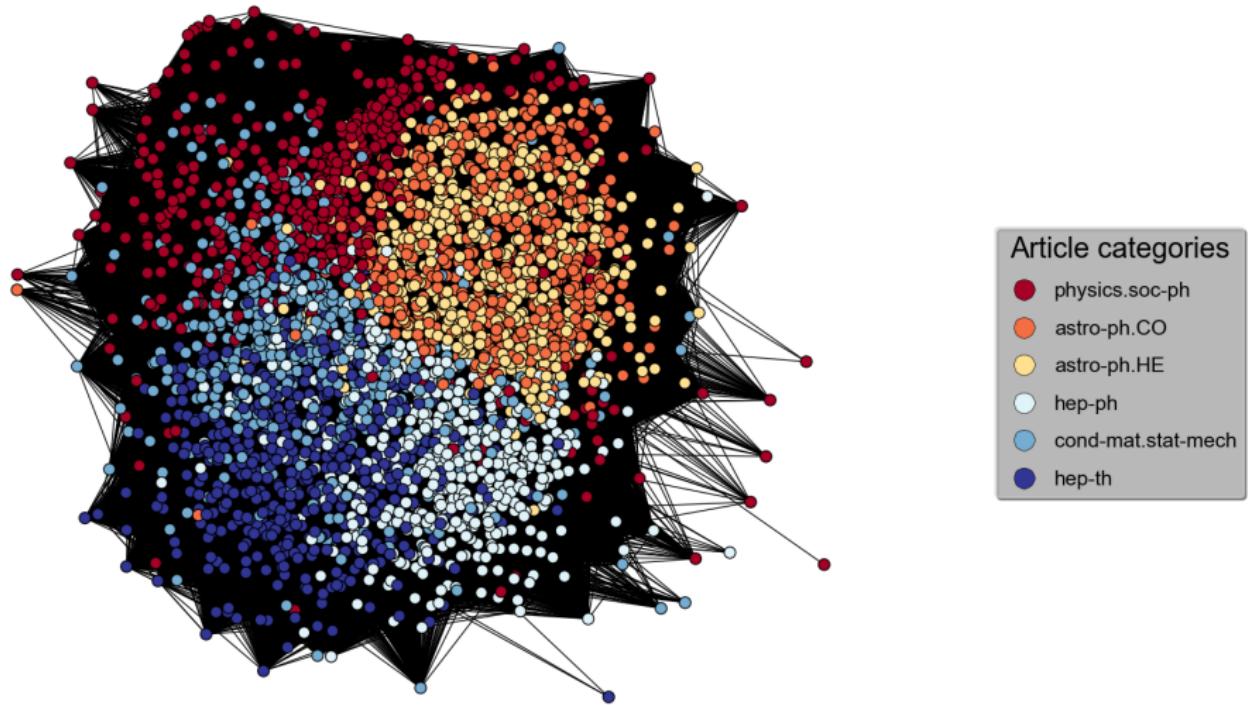
$$w_{\alpha\beta} \in [0, 1],$$

$$\begin{aligned} w_{\alpha\beta} &= \frac{(13 \times 2) + (43 \times 5)}{55.02 \times 34.28} = \\ &= \frac{241}{1886.09} \simeq 0.13. \end{aligned}$$

| | | | | | | | |
|----------|---|----|---|----|---|----|----|
| α | 2 | 43 | 0 | 18 | 0 | 11 | 27 |
|----------|---|----|---|----|---|----|----|

| | | | | | | | |
|---------|----|---|---|---|----|---|---|
| β | 13 | 5 | 9 | 0 | 30 | 0 | 0 |
|---------|----|---|---|---|----|---|---|

$C_1 \quad C_2 \quad C_3 \quad C_4 \quad C_5 \quad C_6 \quad C_7$

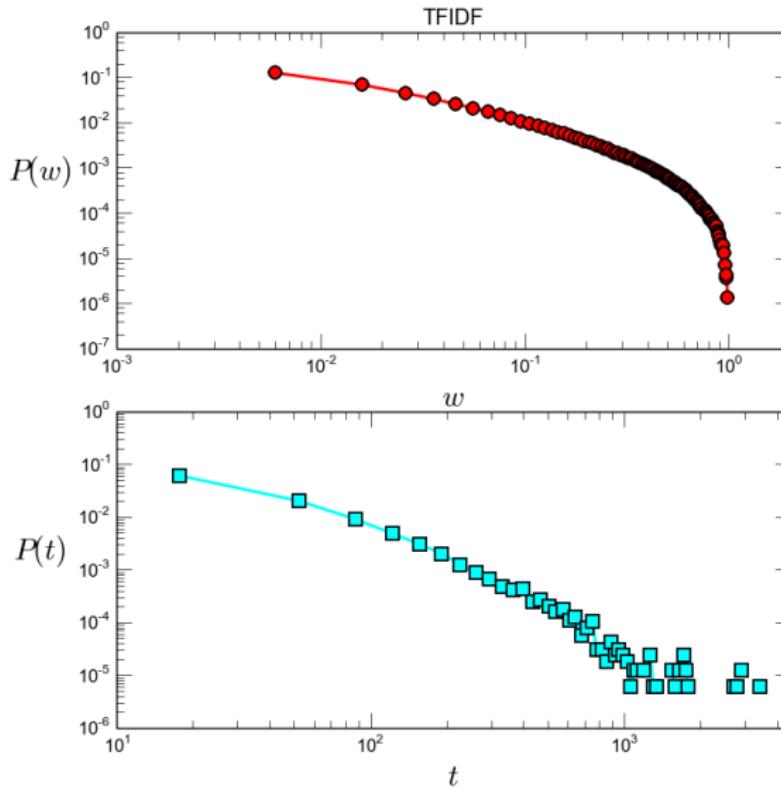


Section 2

Filtering

KUKUAEWm

Available methods



KUKUAEWm

Available methods

PHYSICAL REVIEW E

statistical, nonlinear, and soft matter physics

Highlights Recent Accepted Authors Referees Search About

Information filtering in complex weighted networks

Filippo Radicchi, José J. Ramasco, and Santo Fortunato
Phys. Rev. E **83**, 046101 – Published 1 April 2011

Article

References

Citing Articles (8)

PDF

HTML

Export Citation



ABSTRACT

Many systems in nature, society, and technology can be described as networks, where the vertices are the system's elements, and edges between vertices indicate the interactions between the corresponding elements. Edges may be weighted if the interaction strength is measurable. However, the full network information is often redundant because tools and techniques from network analysis

- Serrano M.A., et al. *Extracting the multiscale backbone of complex weighted networks*. Proc. Natl. Acad. Sci.

(USA) **106** 6483 (2009).

Available methods

Institution: EPFL

Proceedings of the National Academy of Sciences of the United States of America

CURRENT ISSUE // ARCHIVE // NEWS & MULTIMEDIA // FOR AUTHORS // ABOUT PNAS // COLLECTED ARTICLES / BROWSE BY TOPIC / EARLY EDITION

Home > Current Issue > vol. 106 no. 16 > M. Ángeles Serrano, 6483–6488, doi:10.1073/pnas.0808904106

 CrossMark
click for updates

Extracting the multiscale backbone of complex weighted networks

M. Ángeles Serrano^{a,1}, Marián Boguña^b and Alessandro Vespignani^{c,d}

Author Affiliations

Edited by Peter J. Bickel, University of California, Berkeley, CA, and approved March 2, 2009 (received for review September 9, 2008)

Abstract Full Text Authors & Info Figures SI Metrics Related Content   +SI

Abstract

This Issue

April 21, 2009
vol. 106 no. 16
Masthead (PDF)
Table of Contents



◀ PREV ARTICLE ▶ NEXT ARTICLE ▶

 View this article with
LENS beta

Don't Miss

- Radicchi, F., et al. *Information filtering in complex weighted networks*. Physical Review E, **83** 046101. (2011).

Relevant concepts



Relevant concepts

Key features

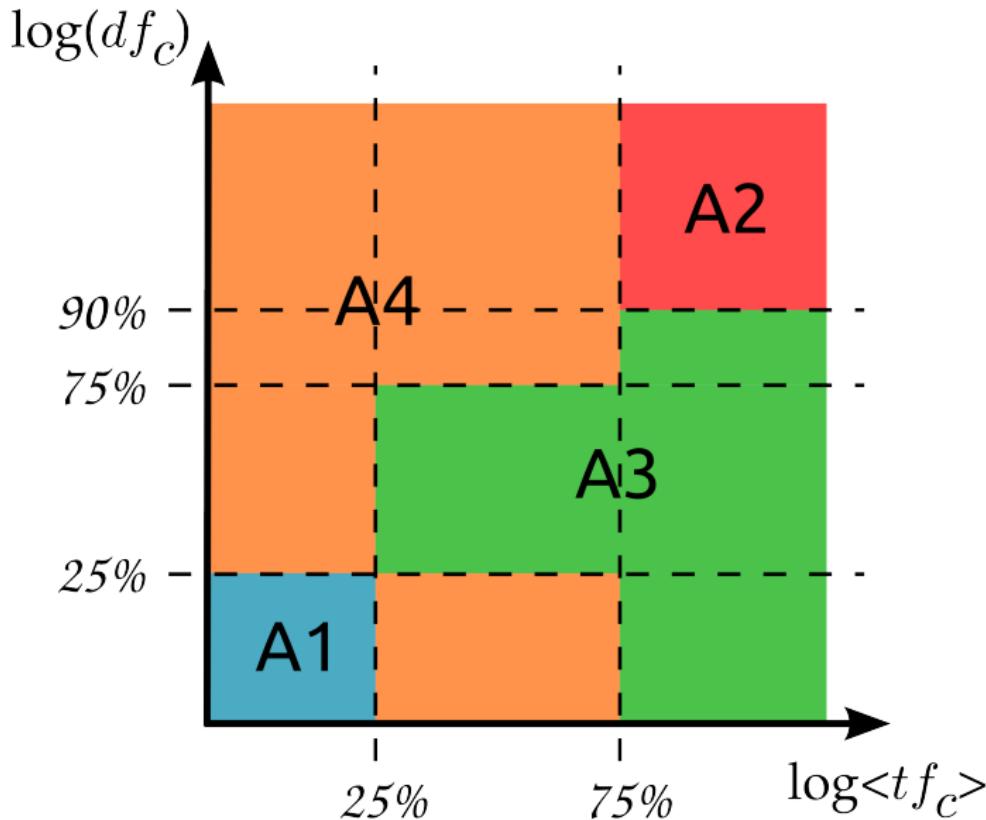
- # of papers a concept appears in

$df_c \rightarrow$ document frequency

- average # of times a concept appears inside a paper

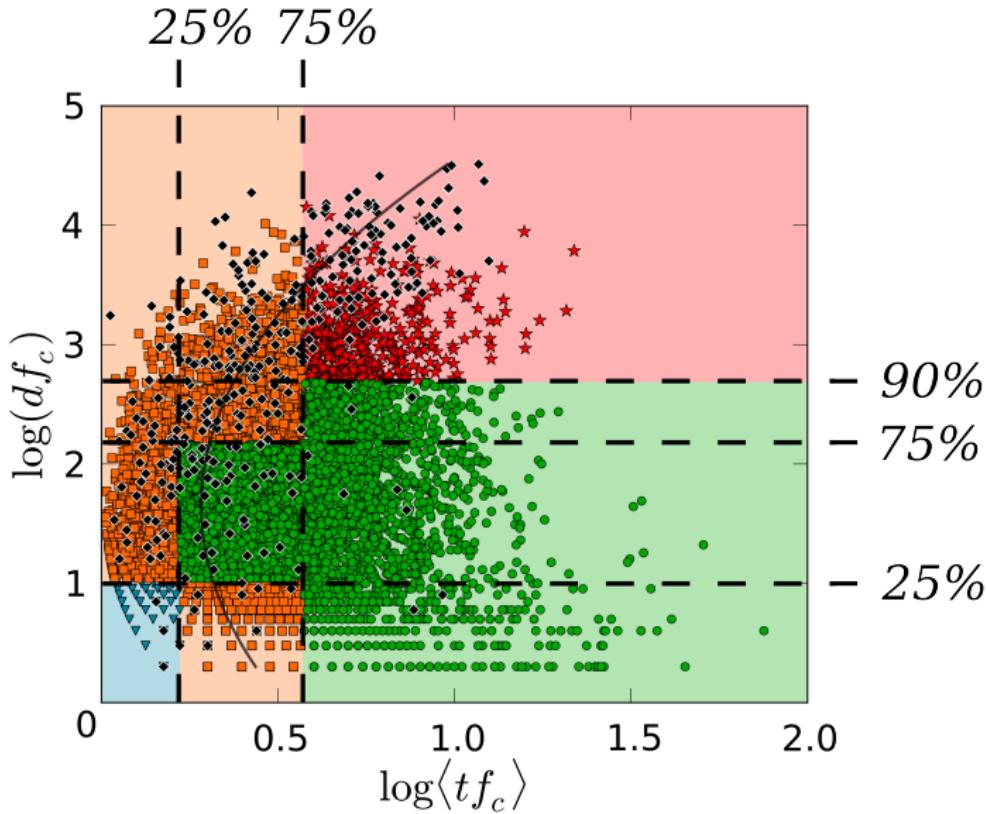
$\langle tf_c \rangle \rightarrow$ average term frequency

Bidimensional tessellation



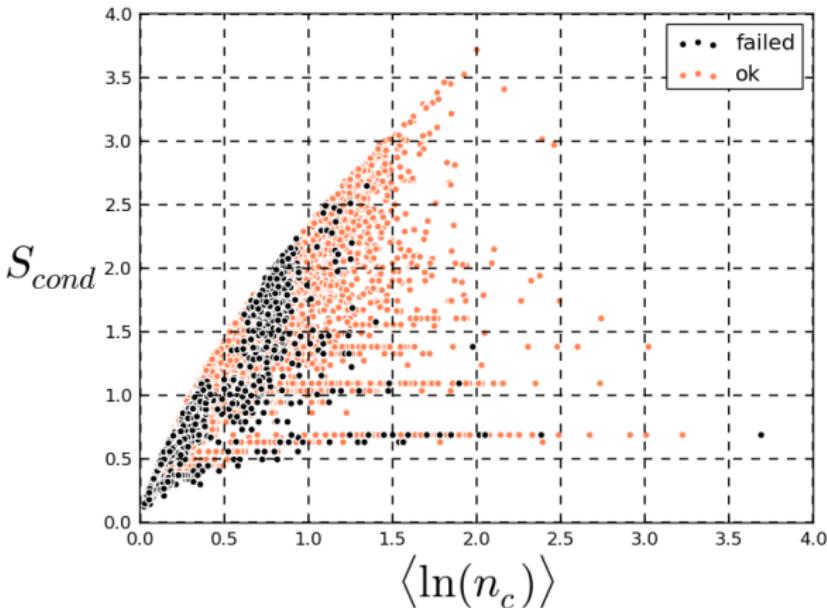
KUKUAEWm

Bidimensional tessellation



KUKUAEWm

Maximum entropy



$$S = - \sum_{j=0}^{\infty} p_c(j) \ln p_c(j)$$

Maximum entropy

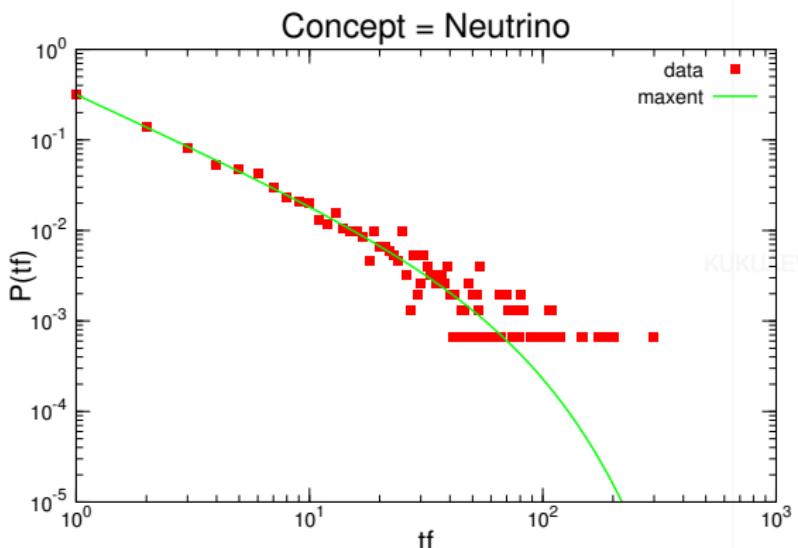
$$\sum_n p_n = 1$$

$$\sum_n p_n n = \langle n \rangle$$

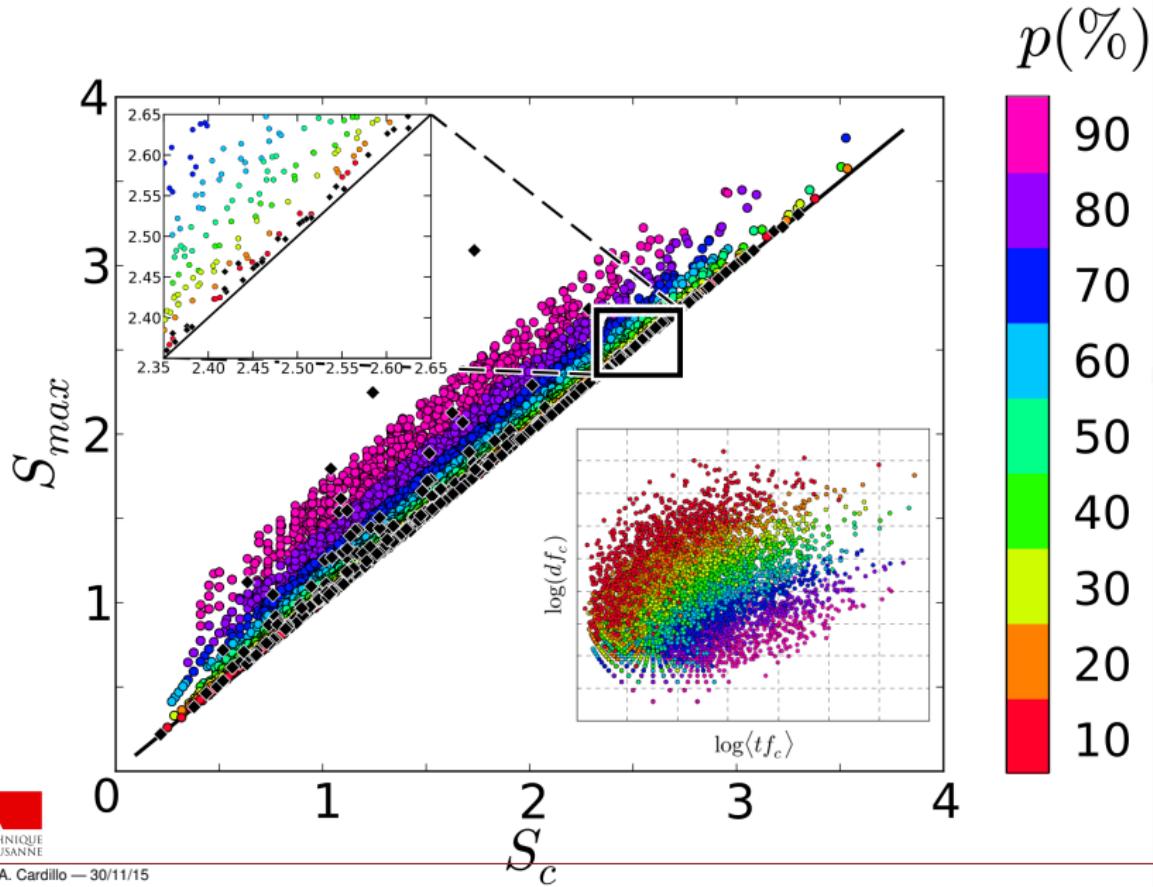
$$\sum_n p_n \ln n = \langle \ln n \rangle$$

$$\ln p_n + \lambda n + \mu \ln n = 0$$

$$p_n = \frac{1}{Z} e^{-\lambda n} n^{-\mu}$$



Maximum entropy



Section 3

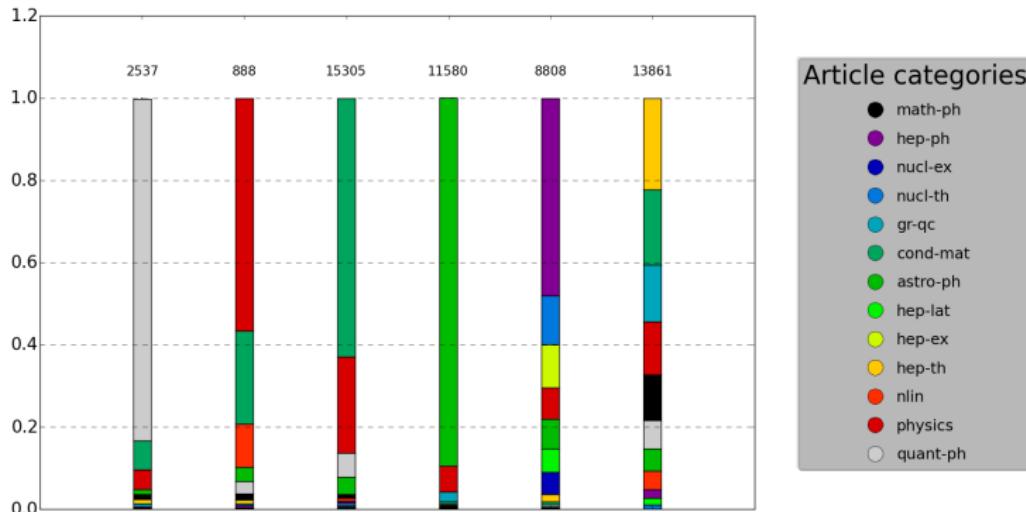
Results

KUKUAEWm

Topological properties

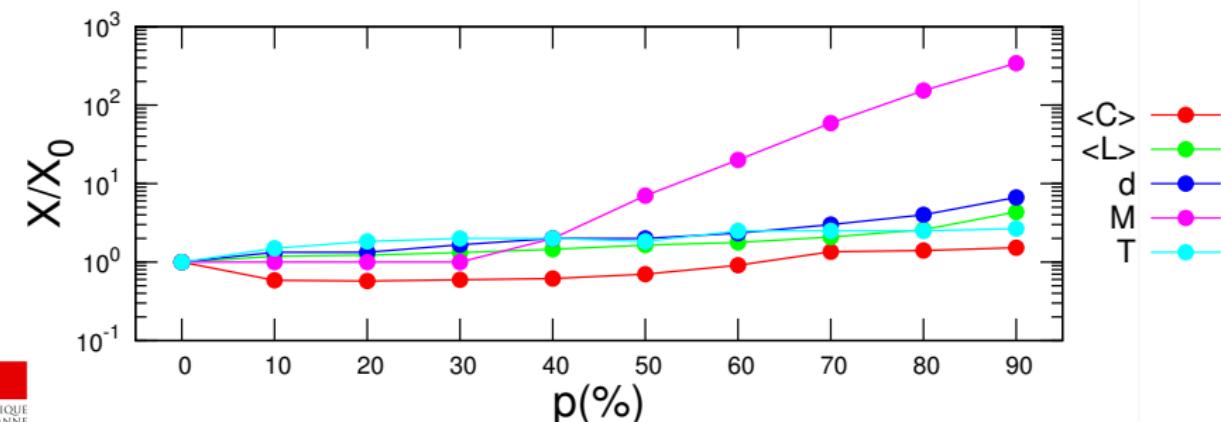
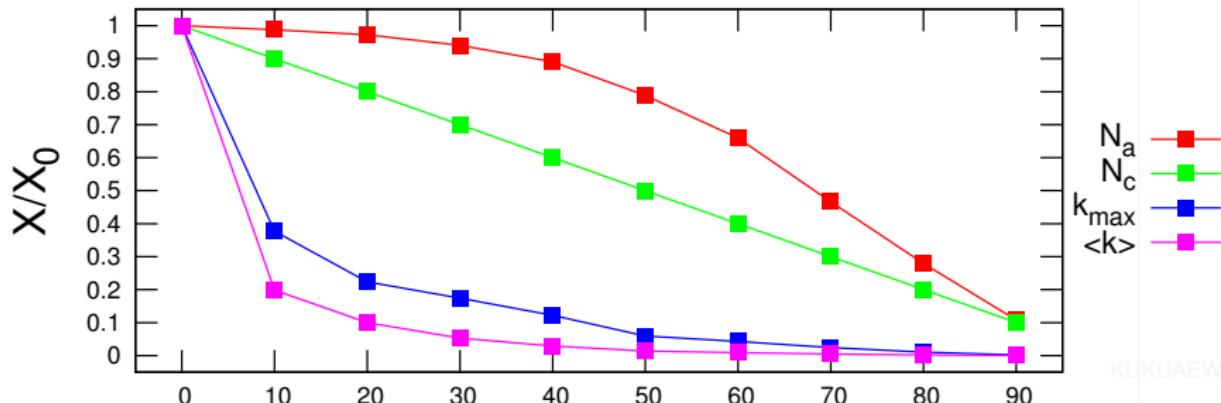
Original network

| N_c | N_a | $\langle k \rangle$ | k_{max} | $\langle C \rangle$ | d | $\langle L \rangle$ | M | T |
|-------|-------|---------------------|-----------|---------------------|-----|---------------------|-----|-----|
| 10661 | 52979 | 19333.522 | 46504 | 0.557 | 3 | 1.635 | 1 | 6 |

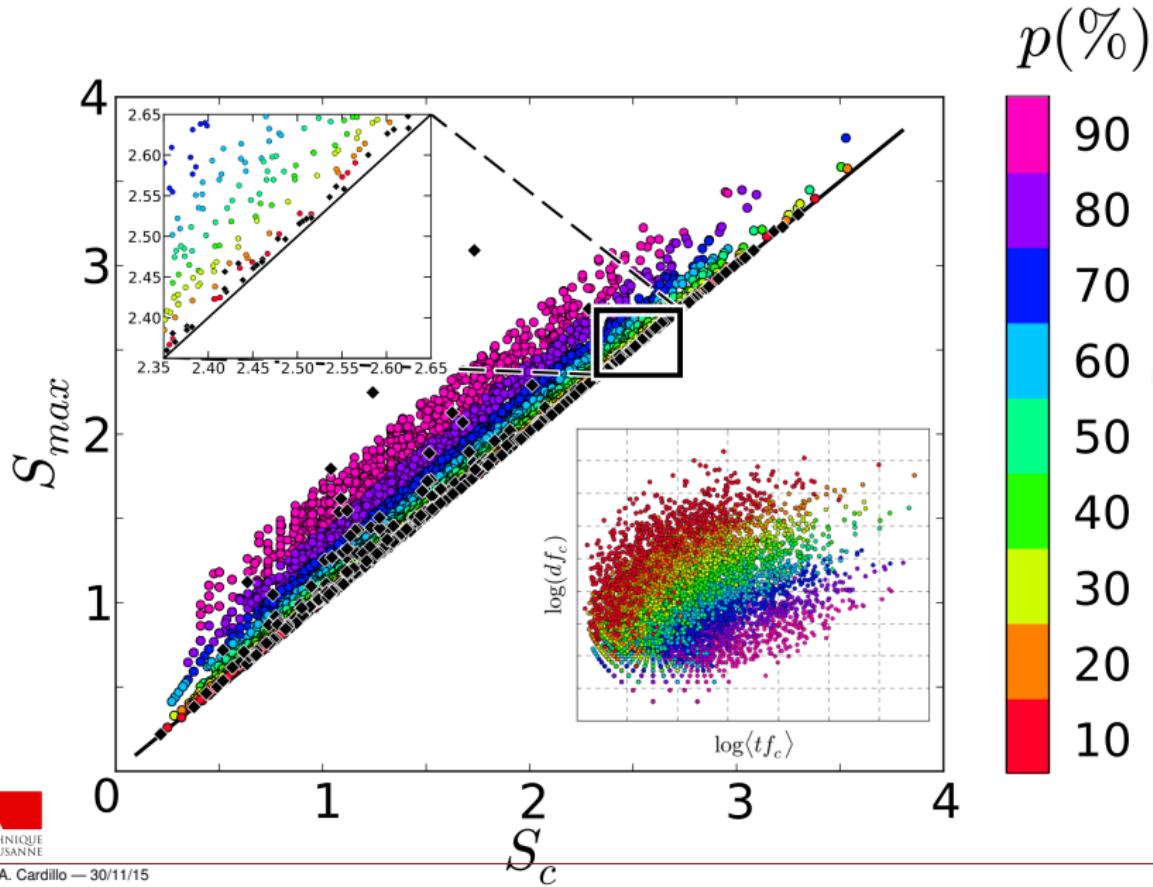


KUKUAEWm

Topological properties



Generic concepts



Community detection

arXiv.org > astro-ph > arXiv:1301.0319

Search or Article-id

Astrophysics > Solar and Stellar Astrophysics

Modules for Experiments in Stellar Astrophysics (MESA): Giant Planets, Oscillations, Rotation, and Massive Stars

Bill Paxton, Matteo Cantiello, Phil Arras, Lars Bildsten, Edward F. Brown, Aaron Dotter, Christopher Mankovich, M. H. Montgomery, Dennis Stello, F. X. Timmes, Richard Townsend

(Submitted on 2 Jan 2013 (v1), last revised 12 Jun 2013 (this version, v2))

We substantially update the capabilities of the open source software package Modules for Experiments in Stellar Astrophysics (MESA), and its one-dimensional stellar evolution module, MESA Star. Improvements in MESA Star's ability to model the evolution of giant planets now extends its applicability down to masses as low as one-tenth that of Jupiter. The dramatic improvement in asteroseismology enabled by the space-based Kepler and CoRoT missions motivates our full coupling of the ADIPLS adiabatic pulsation code with MESA Star. This also motivates a numerical recasting of the Ledoux criterion that is more easily implemented when many nuclei are present at non-negligible abundances. This impacts the way in which MESA Star calculates semi-convective and thermohaline mixing. We exhibit the evolution of 3-8 Msun stars through the end of core He burning, the onset of He thermal pulses, and arrival on the white dwarf cooling sequence. We implement diffusion of angular momentum and chemical abundance that enable calculations of rotating star models which we compare thoroughly with earlier work. We introduce a new treatment of radiation-dominated

Community detection

arXiv.org > astro-ph > arXiv:1306.2314

Search or Article-id

Astrophysics > Cosmology and Nongalactic Astrophysics

Warm Dark Matter as a solution to the small scale crisis: new constraints from high redshift Lyman-alpha forest data

M. Viel, G.D. Becker, J.S. Bolton, M.G. Haehnelt

(Submitted on 10 Jun 2013 ([v1](#)), last revised 26 Aug 2013 (this version, v2))

We present updated constraints on the free-streaming of warm dark matter (WDM) particles derived from an analysis of the Ly α flux power spectrum measured from high-resolution spectra of 25 $z > 4$ quasars obtained with the Keck High Resolution Echelle Spectrometer (HIRES) and the Magellan Inamori Kyocera Echelle (MIKE) spectrograph. We utilize a new suite of high-resolution hydrodynamical simulations that explore WDM masses of 1, 2 and 4 keV (assuming the WDM consists of thermal relics), along with different physically motivated thermal histories. We carefully address different sources of systematic error that may affect our final results and perform an analysis of the Ly α flux power with conservative error estimates. By using a method that samples the multi-dimensional astrophysical and cosmological parameter space, we obtain a lower limit $m_{\text{wdm}} > 3.3 \text{ keV}$ (2sigma) for warm dark matter particles in the form of early decoupled thermal relics. Adding the Sloan Digital Sky Survey (SDSS) Ly α flux power spectrum does not improve this limit. Thermal relics of masses 1 keV, 2 keV and 2.5 keV are disfavoured by the data at about the 9sigma, 4sigma and 3sigma

Community detection

| p | 10 | 20 | 30 | 40 | 50 |
|-----------|--------------|------------------------------|------------------------|-----------------------|------------------------|
| Article | | | | | |
| 1301.0319 | Galaxy | Planet | Planet | White dwarf | Andromeda galaxy |
| | Accretion | Eccentricity | Eclipses | Initial mass function | Large Magellanic Cloud |
| | Stellar mass | Jupiter | Albedo | Globular cluster | Dark matter subhalo |
| | Extinction | Giant planet | Natural satellite | Eclipses | Maser |
| | Supernova | Magneto hydrodynamics | Terrestrial planet | Open cluster | Planetary nebula |
| 1306.2314 | Galaxy | Supernova | Active Galactic Nuclei | White dwarf | Andromeda galaxy |
| | Accretion | Spectral energy distribution | Quasar | Initial mass function | Large Magellanic Cloud |
| | Stellar mass | Neutron star | Pulsar | Globular cluster | Dark matter subhalo |
| | Extinction | Star formation rate | Point spread function | Eclipses | Maser |
| | Supernova | Active Galactic Nuclei | White dwarf | Open cluster | Planetary nebula |

Section 4

Conclusions

KUKUAEWm

Summing up . . .

Take home messages

- We have used the maximum entropy principle to build a method to filter networks of similarity between documents.

IAEWm

Summing up . . .

Take home messages

- We have used the maximum entropy principle to build a method to filter networks of similarity between documents.
- The removal of common concepts allows to retrieve a more well defined structure of documents into topics.

Summing up . . .

Take home messages

- We have used the maximum entropy principle to build a method to filter networks of similarity between documents.
- The removal of common concepts allows to retrieve a more well defined structure of documents into topics.
- The method allows to identify collection dependent “*relevant concepts*” without requiring user validation.

Summing up . . .

What's next? Open questions

- Apply the methodology recursively.

IAEWm

Summing up . . .

What's next? Open questions

- Apply the methodology recursively.
- Validate the method applying it on collection of documents different from scientific papers.

IAEWm

Summing up ...

Reference

A. Martini *et al.* , *Automatic selection of relevant concepts in scientific publications* – in preparation

<http://bifi.es/~cardillo/>

alessio.cardillo@epfl.ch